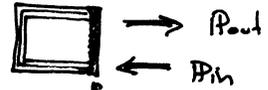


# Radiation problem set (Unified eng.)

The energy balance (heat power eqn.) is:

$$mC \frac{dT}{dt} = P_{in} - P_{out}$$



In steady state:  $P_{in} = P_{out}$

Adiabatic system everywhere, except in the radiator

\* The input power has several parts:

- Dissipated input power:  $P_D = 100 \text{ W}$

- Solar radiation (from albedo):  $P_{S.R.} = S_f a A_r \alpha$

where  $S_f$  is the solar flux:

$$S_f = \sigma T_{sun}^4 \left( \frac{0.695}{149.6} \right)^2 = 1384 \text{ W/m}^2$$

-  $a$  is the Earth's solar albedo,  $a = 0.3$

-  $A_r$  is the radiator area

-  $\alpha$  is the radiator coefficient of absorption

- Direct thermal radiation from the Earth:  $P_{T.R.} = \epsilon \sigma_{EARTH} T_E^4 A_r \left( \frac{r_E}{r_{SAT}} \right)^2 \alpha$

\* Output power is through radiation only:  $P_{out} = \epsilon_r \sigma A_r T_r^4$

Solve for  $T_r \Rightarrow \epsilon_r \sigma A_r T_r^4 = P_D + S_f a A_r \alpha + \epsilon_{EARTH} \sigma T_E^4 A_r \left( \frac{r_E}{r_{SAT}} \right)^2 \alpha$

then  $T_r^4 = \frac{P_D}{\epsilon_r \sigma A_r} + S_f \frac{a}{\sigma} \left( \frac{\alpha}{\epsilon_r} \right) + \left( \frac{\epsilon_{EARTH}}{\epsilon_r} \right) T_E^4 \left( \frac{r_E}{r_{SAT}} \right)^2 \alpha$

$$= \frac{100 \text{ W}}{(0.1)(5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \cdot \text{K}^4})(1 \text{ m}^2)} + \frac{1384 \frac{\text{W}}{\text{m}^2} (0.3)}{(5.67 \times 10^{-8})} + \left( \frac{0.5}{0.1} \right) (300^4) \left( \frac{6378 \text{ km}}{6378 + 500 \text{ km}} \right)^2 0.1$$

then

$$T_r = \boxed{410.68 \text{ K}}$$

(Without looking to "hot" spots

$$T_r = \boxed{369.4 \text{ K}}$$